

Self-similar solutions to Burgers and Navier-Stokes equations

23 August 2018, Banff

Koji Ohkitani (University of Sheffield)

Keywords: scale-invariance, self-similarity, turbulence

Funding: EPSRC

Motivation

No self-similar solutions of Navier-Stokes equations blow up.
Nevertheless it is of interest to see they actually behave.

Cannone-Planchon(1996) proved the existence of forward self-similar solutions for small data by using a fixed-point theorem. For recent developments, see e.g. Chae-Wolf (2018), Bradshaw-Tsai(2018).

We will try to construct asymptotic solutions, using $\nabla \times \omega$.

One of the principal objects of theoretical research is to find the point of view from which the subject appears in the greatest simplicity.

(W. Gibbs)

Plan

0. Introduction
 1. 1D Burgers equation
 2. Examples
 3. Navier-Stokes equations and Hopf equations
 4. Self-similar solutions of Navier-Stokes equations
 5. Summary and outlook
- (with Riccardo Vanon, PDRA)

0. Introduction

Theorem 2 Cannone-Planchon (1996) *Let $u_0 \in \dot{B}_{3,\infty}^0$ be a divergence free vector field that is homogeneous of degree -1 . There exists an universal constant $\eta > 0$ and $T > 0$ such that, if*

$$\|u_0\|_{\dot{B}_{3,\infty}^0} < \eta,$$

then there exists a unique global solution $u(x, t)$ of the Navier-Stokes equations such that

$$u(x, t) = \frac{1}{\sqrt{t}} U \left(\frac{x}{\sqrt{t}} \right)$$

with $U \in \dot{B}_{3,\infty}^0$ and

$$U = S(1)u_0 + W$$

*where $W \in L^3$ satisfies $\|W\|_{L^3} < C(\eta)$.
($S(t)$ = heat operator)*

Initial data for self-similar solution (3D)

$$\lambda u_0(\lambda x) = u_0(x)$$

$$u(x, t) = \frac{1}{\sqrt{t}} U \left(\frac{x}{\sqrt{t}} \right)$$

e.g. $U(x) = \frac{1}{1+|x|}$, $u(x, t) = \frac{1}{|x|+\sqrt{t}} \rightarrow \frac{1}{|x|}$ as $t \rightarrow 0$

Construction of solutions in Besov spaces:

$\alpha > -d$ for homogeneous function of degree α

$$u \sim \frac{1}{x} \text{ in 1D, } \omega \sim \frac{1}{x^2} \text{ in 2D are not included.}$$

cf. $\lambda^d \delta(\lambda x) = \delta(x)$

$$\frac{\partial \mathbf{v}}{\partial t} + (\mathbf{v} \cdot \nabla) \mathbf{v} = \nu \Delta \mathbf{v}, \quad \mathbf{v}(\mathbf{x}, t) = -2\nu \nabla \log \phi$$

$$\mathbf{v}(\mathbf{x}, t) = \frac{\int_{\mathbb{R}^3} \frac{\mathbf{x} - \mathbf{y}}{t} \phi_0(\mathbf{y}) \exp\left(-\frac{|\mathbf{x} - \mathbf{y}|^2}{4\nu t}\right) d\mathbf{y}}{\int_{\mathbb{R}^3} \phi_0(\mathbf{y}) \exp\left(-\frac{|\mathbf{x} - \mathbf{y}|^2}{4\nu t}\right) d\mathbf{y}}$$

A few points deserve emphasis. First, this is only a particular solution to the equation. Second, even so, it is quite complicated and would still require considerable computer analysis to produce a flow field. Third, it is an irrotational solution, as mentioned.

“The Ceaseless Wind: An Introduction to the Theory of Atmospheric Motion,” J.A. Dutton

2. 1D Burgers equation

$$\frac{\partial u}{\partial t} + u \frac{\partial u}{\partial x} = \nu \frac{\partial^2 u}{\partial x^2}$$

static scale invariance

$$x \rightarrow \lambda x, t \rightarrow \lambda^2 t, u \rightarrow \lambda^{-1} u$$

Property 1: if $u(x, t)$ is a solution, so is $\lambda u(\lambda x, \lambda^2 t)$.

Simpler critical case with ϕ where $u = \partial_x \phi$

Property 1': if $\phi(x, t)$ is a solution, so is $\phi(\lambda x, \lambda^2 t)$.

Note velocity potential $[\phi] = [\nu] = L^2/T$

Dynamic scaling $u(x, t) = \frac{1}{\sqrt{2at}} U(\xi, \tau),$

$$\xi = \frac{x}{\sqrt{2at}}, \quad \tau = \frac{1}{2a} \log(1 + 2at)$$

$$\frac{\partial U}{\partial \tau} + U \frac{\partial U}{\partial \xi} = a \frac{\partial}{\partial \xi} (\xi U) + \nu \frac{\partial^2 U}{\partial \xi^2} \quad (\text{Fokker-Planck operator})$$

$$\frac{\partial \Phi}{\partial \tau} + \frac{1}{2} \left(\frac{\partial \Phi}{\partial \xi} \right)^2 = a \xi \frac{\partial \Phi}{\partial \xi} + \nu \frac{\partial^2 \Phi}{\partial \xi^2} \quad (\text{Ornstein-Uhlenbeck op.})$$

quasi-invariance

Steady solution (i.e. self-similar solution)
“source solution”

$$U(\xi) = \frac{U(0) \exp\left(-\frac{a\xi^2}{2\nu}\right)}{1 - \frac{U(0)}{2\nu} \int_0^\xi \exp\left(-\frac{a\eta^2}{2\nu}\right) d\eta}$$

valid only for small data $U(0) < \sqrt{8a\nu/\pi}$

If $u_0 \in L^1$, for $1 \leq p \leq \infty$,

$$t^{\frac{1}{2}\left(1-\frac{1}{p}\right)} \left\| u(x, t) - \frac{1}{\sqrt{2at}} U(\xi) \right\|_p \rightarrow 0 \text{ as } t \rightarrow \infty,$$

where $\xi = \frac{x}{\sqrt{2at}}$

$$\lim_{t \rightarrow 0} \frac{1}{\sqrt{2at}} U(\cdot) = M \delta(x), \quad M \equiv \int_{\mathbb{R}^1} u_0(x) dx$$

e.g. Escobedo-Zuazua(1991), Karch(1999), Biler et al.(2001)

$$\text{cf. } U(0) = \sqrt{\frac{8a\nu}{\pi}} \tanh \frac{M}{4\nu} \approx \sqrt{\frac{a}{2\pi\nu}} M \quad (\nu \gg 1)$$

Three methods of solutions

(1) Simplest method (no linearisation)

$$\frac{U^2}{2} = a\xi U + \nu \frac{dU}{d\xi} = \nu \exp\left(-\frac{a\xi^2}{2\nu}\right) \frac{d}{d\xi} \underbrace{\left(U \exp\left(\frac{a\xi^2}{2\nu}\right)\right)}_{\equiv \tilde{U}}$$

$$2\nu \exp\left(\frac{a\xi^2}{2\nu}\right) \frac{d}{d\xi} \tilde{U} = \tilde{U}^2$$

$$\frac{d\tilde{U}}{d\eta} = \tilde{U}^2, \quad \text{by } \eta \equiv \frac{1}{2\nu} \int_0^\xi \exp\left(-\frac{a\zeta^2}{2\nu}\right) d\zeta$$

(2) Another method (Bernoulli equation)

$$\frac{dU}{d\xi} = \frac{1}{2\nu}U^2 - \frac{a}{\nu}\xi U$$

Linearisation by $V = 1/U$

$$\frac{dV}{d\xi} = \frac{-1}{2\nu}(1 - 2a\xi V)$$

(3) Yet another (awkward) method

Observe

$$U(\xi) = \frac{U(0)\exp\left(-\frac{a\xi^2}{2\nu}\right)}{1 - \frac{U(0)}{2\nu} \int_0^\xi \exp\left(-\frac{a\eta^2}{2\nu}\right) d\eta}$$
$$= U(0)\exp\left(-\frac{a\xi^2}{2\nu}\right) \sum_{n=1}^{\infty} \left\{ \frac{U(0)}{2\nu} \int_0^\xi \exp\left(-\frac{a\eta^2}{2\nu}\right) d\eta \right\}^n$$

$$\text{cf. } \frac{1}{1-r} = 1 + r + r^2 + \dots, \quad |r| < 1$$

Introduce a formal series expansion

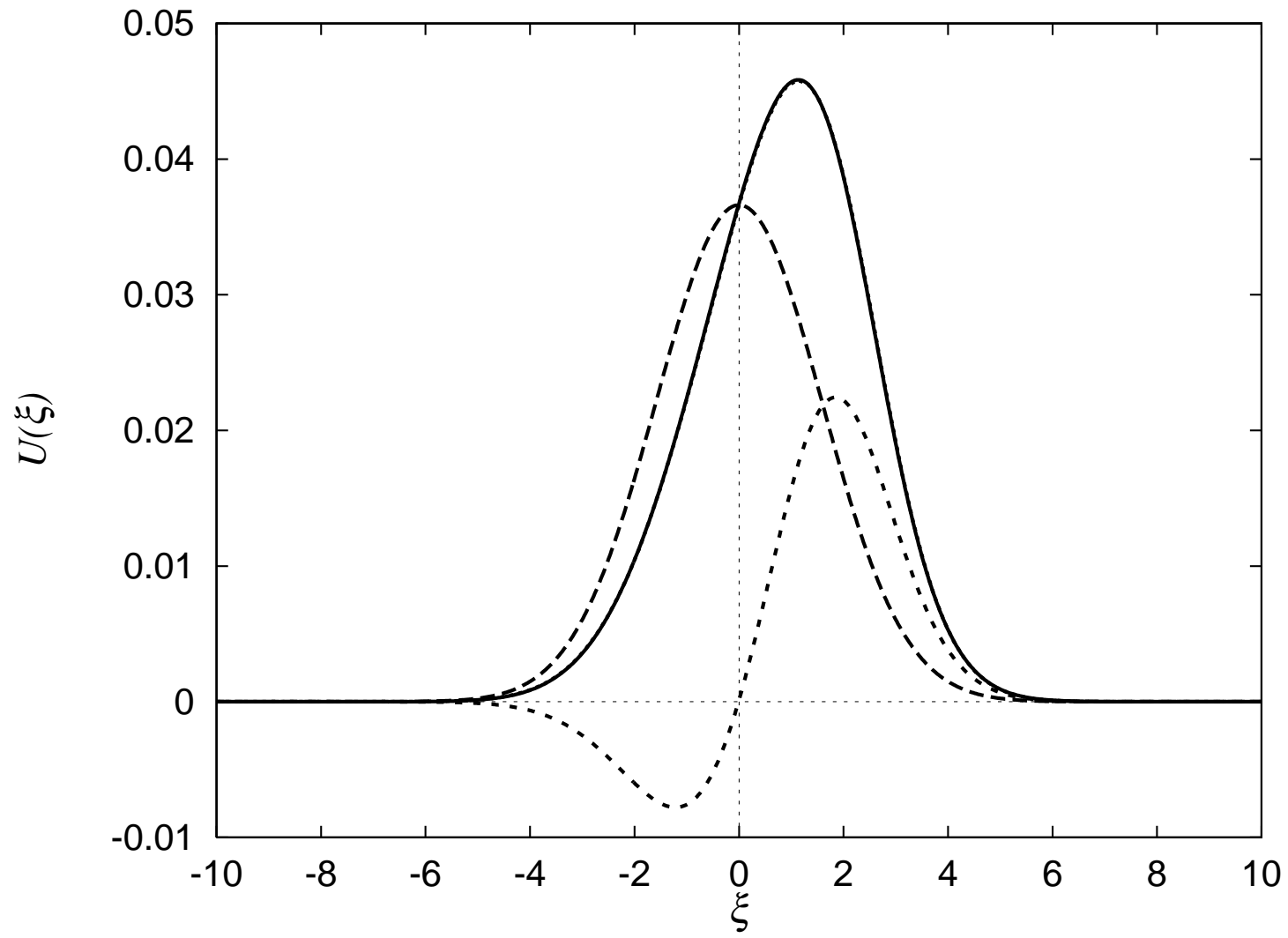
$$\tilde{U} = \epsilon U_1 + \epsilon^2 U_2 + \dots \quad \text{in} \quad \frac{d\tilde{U}}{d\xi} = \frac{1}{2\nu} \exp\left(-\frac{a\xi^2}{2\nu}\right) \tilde{U}^2$$

$$O(\epsilon) : \quad \frac{dU_1}{d\xi} = 0 \rightarrow U_1 = U(0).$$

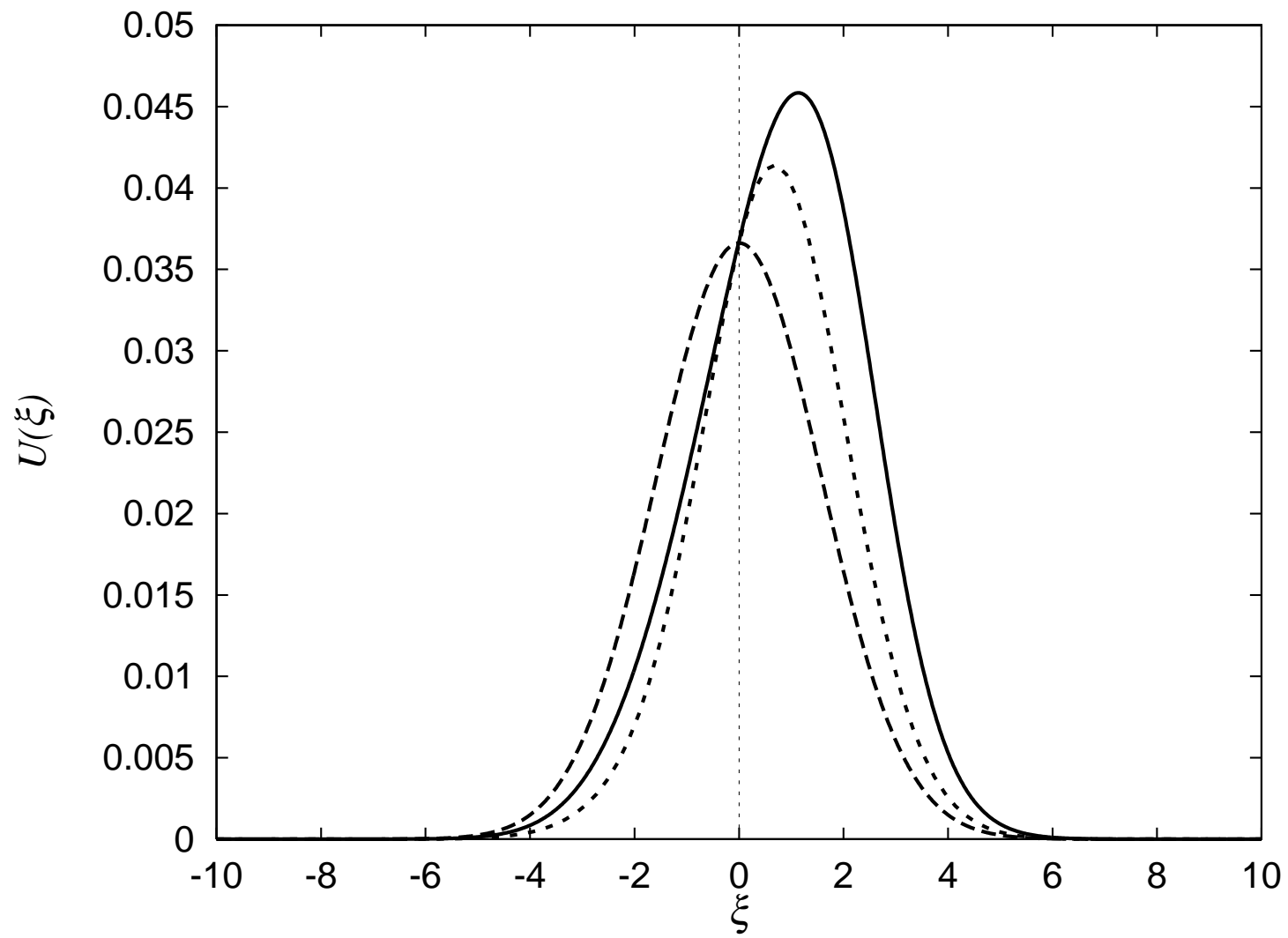
$$O(\epsilon^2) : \quad \frac{dU_2}{d\xi} = \frac{1}{2\nu} \exp\left(-\frac{a\xi^2}{2\nu}\right) U_1^2$$

$$\rightarrow U_2 = \frac{U(0)^2}{2\nu} \int_0^\xi \exp\left(-\frac{a\eta^2}{2\nu}\right) d\eta$$

$$O(\epsilon^3) : \quad \rightarrow U_3 = \frac{U(0)^3}{4\nu^2} \left\{ \int_0^\xi \exp\left(-\frac{a\eta^2}{2\nu}\right) d\eta \right\}^2$$



$\nu=0.05,$
 $a=0.02$



$$\begin{aligned}
 U(\xi) &= \frac{U(0)\exp\left(-\frac{a\xi^2}{2\nu}\right)}{1 - \frac{U(0)}{2\nu} \int_0^\xi \exp\left(-\frac{a\eta^2}{2\nu}\right) d\eta} \\
 &= -2\nu \frac{\partial}{\partial \xi} \log \left\{ U(0)^{-1} - \frac{1}{2\nu} \int_0^\xi \exp\left(-\frac{a\eta^2}{2\nu}\right) d\eta \right\}
 \end{aligned}$$

step 1: find a self-similar solution

step 2: make a substitution
a self-similar heat flow

$$u(x, t) = \frac{1}{\sqrt{2at}} \frac{U(0) \exp\left(-\frac{x^2}{4\nu t}\right)}{1 - \frac{U(0)}{2\nu} \int_0^{\frac{x}{\sqrt{2at}}} \exp\left(-\frac{\eta^2}{4\nu t}\right) d\eta}$$

\implies a more general heat flow

$$u(x, t) = \frac{\int_{-\infty}^{\infty} \frac{x-y}{t} \psi_0(y) \exp\left(-\frac{(x-y)^2}{4\nu t}\right) dy}{\int_{-\infty}^{\infty} \psi_0(y) \exp\left(-\frac{(x-y)^2}{4\nu t}\right) dy}$$

2. Examples

Burgers equations in several dimensions

1D

$$U(\xi) = \frac{U(0) \exp\left(-\frac{a\xi^2}{2\nu}\right)}{1 - R}$$

$$R = \frac{U(0)}{2\nu} \int_0^\xi \exp\left(-\frac{a\xi'^2}{2\nu}\right) d\xi'$$

or

$$U(\xi) = -2\nu \frac{\partial}{\partial \xi} \log(1 - R)$$

2D

$$\frac{\partial U_1}{\partial \xi_2} = \frac{\frac{\partial U_1}{\partial \xi_2}(0) \exp\left(-\frac{a(\xi_1^2 + \xi_2^2)}{2\nu}\right)}{(1 - R)^2}$$

$$R = \frac{\frac{\partial U_1}{\partial \xi_2}(0)}{2\nu} \int_0^{\xi_1} \exp\left(-\frac{a\xi^2}{2\nu}\right) d\xi \int_0^{\xi_2} \exp\left(-\frac{a\eta^2}{2\nu}\right) d\eta$$

or

$$\frac{\partial U_1}{\partial \xi_2} = -2\nu \frac{\partial^2}{\partial \xi_1 \partial \xi_2} \log(1 - R)$$

3D

$$\frac{\partial^2 U_1}{\partial \xi_2 \partial \xi_3} = \frac{\partial^2 U_1}{\partial \xi_2 \partial \xi_3}(0) \exp\left(-\frac{a(\xi_1^2 + \xi_2^2 + \xi_3^2)}{2\nu}\right) \frac{1 - R}{(1 - R)^3}$$

$$R = \frac{\frac{\partial^2 U_1}{\partial \xi_2 \partial \xi_3}(0)}{2\nu} \int_0^{\xi_1} \exp\left(-\frac{a\xi^2}{2\nu}\right) d\xi \int_0^{\xi_2} \exp\left(-\frac{a\eta^2}{2\nu}\right) d\eta \int_0^{\xi_3} \exp\left(-\frac{a\zeta^2}{2\nu}\right) d\zeta$$

or

$$\frac{\partial^2 U_1}{\partial \xi_2 \partial \xi_3} = -2\nu \frac{\partial^3}{\partial \xi_1 \partial \xi_2 \partial \xi_3} \log(1 - R)$$

Common: $\mathbf{U} = \nabla\phi$, $\phi = -2\nu \log(1 - R)$

2D Navier-Stokes equation

“A Mathematical Model Illustrating the Theory of Turbulence,” (1948)

*Burgers vortex $\mathbf{u} = (u_r, u_\theta, u_\phi) = (-ar, v(r), 2az)$

$$\omega(r) = \frac{a\Gamma}{2\pi\nu} \exp\left(-\frac{ar^2}{2\nu}\right), \quad \Gamma \equiv \int_{\mathbb{R}^2} \omega_0(\mathbf{x}) d\mathbf{x}$$

$$v(r) = \frac{\Gamma}{2\pi r} \left(1 - \exp\left(-\frac{ar^2}{2\nu}\right)\right)$$

*Oseen vortex $\omega(r, t) = \frac{\Gamma}{4\pi\nu t} \exp\left(-\frac{r^2}{4\nu t}\right), \quad \omega(r, 0) = \Gamma\delta(\cdot)$

If $\omega_0 \in L^2$, for $1 \leq p \leq \infty$,

$$t^{1-\frac{1}{p}} \left\| \omega(\mathbf{x}, t) - \frac{1}{2at} \Omega(\boldsymbol{\xi}) \right\|_p \rightarrow 0 \text{ as } t \rightarrow \infty$$

$$\Omega(\boldsymbol{\xi}) = \frac{a\Gamma}{2\pi\nu} \exp\left(-\frac{a|\boldsymbol{\xi}|^2}{2\nu}\right), \quad \boldsymbol{\xi} = \frac{\mathbf{x}}{\sqrt{2at}}$$

$$\frac{1}{2at} \Omega(\boldsymbol{\xi}) = \frac{\Gamma}{4\pi\nu t} \exp\left(-\frac{|\mathbf{x}|^2}{4\nu t}\right)$$

$$\lim_{t \rightarrow 0} \Omega(\cdot) = \Gamma \delta(\mathbf{x}), \quad \Gamma \equiv \int_{\mathbb{R}^2} \omega_0(\mathbf{x}) d\mathbf{x}$$

e.g. Gallay & Wayne(2005)

The below are all 3D Navier-Stokes equations. ($\mathbf{r} \equiv \mathbf{x} - \mathbf{y}$)

$$\frac{\partial \psi}{\partial t} = \frac{3}{4\pi} \int_{\mathbb{R}^3} \frac{\mathbf{r} \times (\nabla \times \psi(\mathbf{y})) \mathbf{r} \cdot (\nabla \times \psi(\mathbf{y}))}{|\mathbf{r}|^5} d\mathbf{y} + \nu \Delta \psi + a(\mathbf{x} \cdot \nabla) \psi$$

$$\frac{\partial \mathbf{u}}{\partial t} + \mathbf{u} \cdot \nabla \mathbf{u} = -\nabla p + \nu \Delta \mathbf{u} + a(\mathbf{x} \cdot \nabla) \mathbf{u} + a\mathbf{u}$$

$$\frac{\partial \boldsymbol{\omega}}{\partial t} + \mathbf{u} \cdot \nabla \boldsymbol{\omega} = \boldsymbol{\omega} \cdot \nabla \mathbf{u} + \nu \Delta \boldsymbol{\omega} + a(\mathbf{x} \cdot \nabla) \boldsymbol{\omega} + 2a\boldsymbol{\omega}$$

$$\frac{\partial \boldsymbol{\chi}}{\partial t} = \nabla \times (\mathbf{u} \times \boldsymbol{\chi} + 2(\boldsymbol{\omega} \cdot \nabla) \mathbf{u}) + \nu \Delta \boldsymbol{\chi} + \underbrace{a(\mathbf{x} \cdot \nabla) \boldsymbol{\chi} + 3a\boldsymbol{\chi}}_{=a\nabla \cdot (\mathbf{x} \otimes \boldsymbol{\chi})}$$

3. Navier-Stokes equations and Hopf equations (interlude)

3D Navier-Stokes equations

$$\frac{\partial \mathbf{u}}{\partial t} + \mathbf{u} \cdot \nabla \mathbf{u} + \nabla p = \nu \Delta \mathbf{u}, \quad \nabla \cdot \mathbf{u} = 0,$$

Property 1: if $\mathbf{u}(\mathbf{x}, t)$ is a solution, so is $\lambda \mathbf{u}(\lambda \mathbf{x}, \lambda^2 t)$.

Dynamic scaling

$$\mathbf{u}(\mathbf{x}, t) = \frac{1}{\sqrt{2at}} \mathbf{U}(\boldsymbol{\xi}, \tau),$$

$$\boldsymbol{\xi} = \frac{\mathbf{x}}{\sqrt{2at}}, \quad \tau = \int_0^t \frac{ds}{\lambda(s)^2} = \frac{1}{2a} \log(1 + 2at),$$

the scaled Navier-Stokes equations (the Leray equations)

$$\frac{\partial \mathbf{U}}{\partial \tau} + \mathbf{U} \cdot \nabla_{\boldsymbol{\xi}} \mathbf{U} = -\nabla_{\boldsymbol{\xi}} P + \nu \Delta_{\boldsymbol{\xi}} \mathbf{U} + a(\boldsymbol{\xi} \cdot \nabla_{\boldsymbol{\xi}} \mathbf{U} + \mathbf{U}), \quad \nabla_{\boldsymbol{\xi}} \cdot \mathbf{U} = 0.$$

Characteristic functional of velocity $\Phi[\boldsymbol{\theta}, t] = \left\langle \exp \left(i \int \mathbf{u}(\mathbf{x}, t) \cdot \boldsymbol{\theta}(\mathbf{x}) d\mathbf{x} \right) \right\rangle$

Hopf equation (1952)

$$\frac{\partial \Phi}{\partial t} = i \int d\mathbf{x} \theta_j^\perp(\mathbf{x}) \frac{\partial}{\partial x_k} \frac{\delta^2 \Phi}{\delta \theta_j(\mathbf{x}) \delta \theta_k(\mathbf{x})} + \nu \int d\mathbf{x} \theta_j(\mathbf{x}) \Delta \frac{\delta \Phi}{\delta \theta_j(\mathbf{x})}, \quad \equiv L\Phi$$

$$\left(\frac{\partial}{\partial t} - L \right) \Phi \left[\lambda^2 \boldsymbol{\theta}(\lambda \mathbf{x}), \lambda^{-2} t \right] = 0, \quad \text{Rosen(1982)}$$

Extends to d -dimensions, hence,

Property 2: if $\Phi[\boldsymbol{\theta}(\mathbf{x}), t]$ is a solution, so is $\Phi[\lambda^{d-1} \boldsymbol{\theta}(\lambda \mathbf{x}), \lambda^{-2} t]$.

For vorticity char. functional, $\Phi[\lambda^{d-2} \boldsymbol{\theta}(\lambda \mathbf{x}), \lambda^{-2} t]$

For vorticity gradient char. functional, $\Phi[\lambda^{d-3} \boldsymbol{\theta}(\lambda \mathbf{x}), \lambda^{-2} t]$

See also Foias(1972,1973) for an alternative formulation based on the Liouville equation.

4. Self-similar solutions of Navier-Stokes equations

3D Navier-Stokes equations written in $\chi(x, t) = \nabla \times \omega$

$$\frac{\partial \chi}{\partial t} = \nabla \times (u \times \chi + 2(\omega \cdot \nabla)u) + \nu \Delta \chi$$

After scaling, $X(\xi, \tau)$ satisfies:

$$\frac{\partial X}{\partial \tau} = \nabla \times (U \times X + 2(\Omega \cdot \nabla)U) + \underbrace{a \nabla \cdot (\xi \otimes X)}_{=3aX + a(\xi \cdot \nabla)X} + \nu \Delta X$$

$$\chi = \frac{X(\xi)}{(2at)^{3/2}}, \quad \omega = \frac{\Omega(\xi)}{2at}, \quad u = \frac{U(\xi)}{(2at)^{1/2}}$$

$$\begin{cases} \mathbf{X}(\xi) = \epsilon \mathbf{X}_1(\xi) + \epsilon^2 \mathbf{X}_2(\xi) + \dots \\ U(\xi) = \epsilon U_1(\xi) + \epsilon^2 U_2(\xi) + \dots \\ \Omega(\xi) = \epsilon \Omega_1(\xi) + \epsilon^2 \Omega_2(\xi) + \dots \end{cases}$$

Homogeneous Fokker-Planck equation

$$O(\epsilon) : \quad \Delta \mathbf{X}_1 + \frac{a}{\nu} \nabla \cdot (\xi \otimes \mathbf{X}_1) = 0$$

$$\mathbf{X}_1 = \int \mathbf{X}_0(\eta) G(\xi - \eta) d\eta = \mathbf{X}_0 * G$$

where $G(\xi) = \left(\frac{a}{2\pi\nu}\right)^{3/2} \exp\left(-\frac{a}{2\nu}|\xi|^2\right)$

Inhomogeneous Fokker-Planck equation

$$O(\epsilon^2) : \quad \Delta \mathbf{X}_2 + \frac{a}{\nu} \nabla \cdot (\boldsymbol{\xi} \otimes \mathbf{X}_2) = -\frac{1}{\nu} \nabla \times \underbrace{\{\mathbf{U}_1 \times \mathbf{X}_1 + 2(\boldsymbol{\Omega}_1 \cdot \nabla)\} \mathbf{U}_1}_{\equiv \mathbf{F}}$$

$$\mathbf{U}_1 = \mathbf{U}_0 * G, \quad \boldsymbol{\Omega}_1 = \boldsymbol{\Omega}_0 * G$$

$$\mathbf{X}_2 = -\frac{1}{\nu} g * (\nabla \times \mathbf{F})$$

$$= \frac{-1}{4\pi\nu} \left(\frac{a}{\nu} \exp\left(-\frac{ar^2}{2\nu}\right) \int_0^r \exp\left(\frac{as^2}{2\nu}\right) ds - \frac{1}{r} \right) * \nabla \times \{\mathbf{U}_1 \times \mathbf{X}_1 + 2(\boldsymbol{\Omega}_1 \cdot \nabla)\mathbf{U}_1\}$$

where g is the Green's function for the Fokker-Planck operator

$$g \equiv \frac{1}{4\pi} \left(\frac{a}{\nu} e^{-\frac{ar^2}{2\nu}} \int_0^r e^{\frac{as^2}{2\nu}} ds - \frac{1}{r} \right)$$

$$\mathbf{X}_2 = -\frac{1}{\nu} \underbrace{g^*}_{L^2} \underbrace{\nabla \times}_{L^{-1}} \left\{ \underbrace{U_1}_{L^2} \times \mathbf{X}_1 + 2 \underbrace{(\Omega_1 \cdot \nabla)}_L \underbrace{U_1}_L \right\}$$

cf.

3D Burgers

$$\frac{\partial^2 U_1}{\partial \xi_2 \partial \xi_3} = \frac{\partial^2 U_1}{\partial \xi_2 \partial \xi_3}(0) \exp\left(-\frac{a(\xi_1^2 + \xi_2^2 + \xi_3^2)}{2\nu}\right) \frac{1-R}{(1-R)^3}$$

$$R = \frac{\frac{\partial^2 U_1}{\partial \xi_2 \partial \xi_3}(0)}{2\nu} \underbrace{\int_0^{\xi_1} \exp\left(-\frac{a\xi^2}{2\nu}\right) d\xi}_L \underbrace{\int_0^{\xi_2} \exp\left(-\frac{a\eta^2}{2\nu}\right) d\eta}_L \underbrace{\int_0^{\xi_3} \exp\left(-\frac{a\zeta^2}{2\nu}\right) d\zeta}_L$$

3D Navier-Stokes

$$t^{\frac{3}{2}\left(1-\frac{1}{p}\right)} \left\| \chi(\mathbf{x}, t) - \frac{\mathbf{X}_1(\boldsymbol{\xi})}{(2at)^{3/2}} \right\|_p \rightarrow 0 \text{ as } t \rightarrow \infty$$

$$\text{Note : } \frac{\mathbf{X}_1(\boldsymbol{\xi})}{(2at)^{3/2}} \rightarrow \mathbf{X}_0 * \delta = \mathbf{X}_0(\cdot) \text{ as } t \rightarrow 0,$$

where $\mathbf{X}_0(\boldsymbol{\xi})$ is singular, like the Dirac mass.

Initial data

$$\begin{aligned} X_0 &= P c \delta = (I - \Delta^{-1} \nabla \nabla \cdot) c \delta \\ &= (\delta_{ij} - \Delta^{-1} \partial_i \partial_j) c_j \delta \end{aligned}$$

P =Leray projection, c =const.

$$\begin{aligned} X_1 &= X_0 * G \\ &= (\delta_{ij} - \Delta^{-1} \partial_i \partial_j) c_j \delta * G \\ &= (\delta_{ij} - \Delta^{-1} \partial_i \partial_j) c_j G \end{aligned}$$

$$\frac{\partial^2}{\partial x_i \partial x_j} \Delta^{-1} f = \frac{\delta_{ij}}{3} f + T_{ij}[f]$$

$$T_{ij}[f](\mathbf{x}) \equiv \frac{1}{4\pi} \text{PV} \int \left(\frac{\delta_{ij}}{|\mathbf{x} - \mathbf{y}|^3} - \frac{3(x_i - y_i)(x_j - y_j)}{|\mathbf{x} - \mathbf{y}|^5} \right) f(\mathbf{y}) d\mathbf{y}$$

$$(\mathbf{X}_1)_i = \frac{2}{3} c_i G - T_{ij}[c_j G],$$

$$(\mathbf{X}_1)_i = c_i G + R_i R_j c_j G$$

or,

$$\mathbf{X}_1 = \frac{2}{3} \mathbf{c}G - \frac{1}{4\pi} \text{PV} \int \left(\frac{\mathbf{c}G(\mathbf{y})}{|\mathbf{x} - \mathbf{y}|^3} - \frac{3(\mathbf{x} - \mathbf{y})(\mathbf{x} - \mathbf{y}) \cdot \mathbf{c}G(\mathbf{y})}{|\mathbf{x} - \mathbf{y}|^5} \right) d\mathbf{y}$$

Step 1:

$$\mathbf{X} = \epsilon \mathbf{X}_1 + \epsilon^2 \mathbf{X}_2 + \dots$$

or, component-wise

$$X = \epsilon X_1 \left(1 + \epsilon \frac{X_2}{X_1} + \dots \right)$$

In practice, rationalisation in the spirit of Padé approximation may be useful

$$X \approx \frac{X_1}{1 - \frac{X_2}{X_1}}$$

(for step 2) corresponding general heat flow

$$\chi_1 = \int \chi_0(\mathbf{y}) \hat{G}(\mathbf{x} - \mathbf{y}) d\mathbf{y} = \chi_0 * \hat{G}, \quad \mathbf{u}_1 = \mathbf{u}_0 * \hat{G}, \quad \boldsymbol{\omega}_1 = \boldsymbol{\omega}_0 * \hat{G}$$

$$\hat{G} = \frac{1}{(4\pi\nu t)^{3/2}} \exp\left(-\frac{|\mathbf{x}|^2}{4\nu t}\right)$$

$$\chi_2 = \frac{-1}{4\pi\nu} \left(\frac{a}{\nu} \exp\left(-\frac{ar^2}{2\nu}\right) \int_0^r \exp\left(\frac{as^2}{2\nu}\right) ds - \frac{1}{r} \right) * \nabla \times \{ \mathbf{u}_1 \times \chi_1 + 2(\boldsymbol{\omega}_1 \cdot \nabla) \mathbf{u}_1 \}$$

$$\chi = \epsilon \chi_1 + \epsilon^2 \chi_2 + \dots$$

Component-wise

$$\chi = \epsilon \chi_1 \left(1 + \epsilon \frac{\chi_2}{\chi_1} + \dots \right), \quad \chi \approx \frac{\chi_1}{1 - \frac{\chi_2}{\chi_1}}$$

1D In $U_1 = U_0 * G$, $U_0 = M\delta \Rightarrow U_1 = MG$

$$\frac{U(\xi)}{\sqrt{2at}} \rightarrow M\delta, \quad M = \int u_0.$$

2D In $\Omega_1 = \Omega_0 * G$, $\Omega_0 = \Gamma\delta \Rightarrow \Omega_1 = \Gamma G$

$$\frac{\Omega(\xi)}{2at} \rightarrow \Gamma\delta, \quad \Gamma = \int \omega_0.$$

3D In $X_1 = X_0 * G$, $X_0 = KP\delta c \Rightarrow X_1 = KPGc$

$$\frac{X(\xi)}{(2at)^{3/2}} \rightarrow KP\delta c, \quad Kc = \int \nabla \times \omega_0 = \int d\mathbf{S} \times \omega_0$$

5. Summary and outlook

Self-similar solutions of the Navier-Stokes equations

Two kinds of scale-invariance

d -dimensional Burgers equations

3D Navier-Stokes equations:

asymptotic form of self-similar decaying solutions

Numerical evaluations ? Implications on statistical solutions ?