

Ancient solutions to geometric flows

Panagiota Daskalopoulos

Columbia University

Banff
May 2014

- We will discuss **ancient** or **eternal** solutions to geometric flows, that is solutions that exist for all time

$$-\infty < t < T \quad \text{where } T \in (-\infty, +\infty].$$

- Such solutions appear as **blow up** limits near a singularity.

We will discuss:

- the problem **classification** of ancient or eternal solutions (mean curvature flow, Ricci flow, Yamabe flow), and
- methods of constructing new ancient solutions from the **gluing** of two or more **self-similar** solutions (Yamabe flow).

Ancient and Eternal solutions

- **Definition:** A solution to a parabolic equation is called **ancient** if it is defined for all time $-\infty < t < T$ and **eternal** if it is defined for all time $-\infty < t < +\infty$.
- The **classification** of ancient and eternal solutions often plays a crucial role in understanding the singularities of the flow.
- **Solitons** (self-similar solutions) are examples of ancient or eternal solutions and often models of singularities.
- However, there exists **other** interesting ancient or eternal solutions that are **not solitons** and they often can be visualized as obtained from the gluing of one or more solitons.
- **Definition:** An ancient solution is called of **type I** if:

$$\limsup_{t \rightarrow -\infty} (|t| \max_M R(\cdot, t)) < \infty.$$

Otherwise, it is called of **type II**.

Liouville type Theorem for the heat equation

- Let M be a complete **non-compact** Riemannian manifold of dimension $n \geq 2$ with $\text{Ricci}(M) \geq 0$.
- **Yau - 1975:** Any **positive harmonic** function u on M must be **constant**.
- This is the analogue of Liouville's Theorem for **harmonic** functions on \mathbb{R}^n .
- **Question:** Does the analogue of Yau's theorem hold for **positive** solutions of the heat equation

$$u_t = \Delta u \quad \text{on } M?$$

- **Answer:** No. Example $u(x, t) = e^{x+t}$ on $M := \mathbb{R}^n$.

Liouville type Theorem for the heat equation

- Souplet and Zhang - 2006:

(a) Let u be a **positive ancient** solution to the heat equation on $M \times (-\infty, T)$ such that

$$u(x, t) = e^{o(d(x) + \sqrt{|t|})} \quad \text{as } d(x) \rightarrow \infty.$$

Then u is a **constant**.

(b) Let u be an ancient solution to the heat equation such that

$$u(x, t) = o(d(x) + \sqrt{|t|}) \quad \text{as } d(x) \rightarrow \infty.$$

Then u is a **constant**.

- **Proof:** By using a **local gradient estimate** on large appropriately scaled parabolic cylinders.

The sub-critical semi-linear heat equation

- Consider **nonnegative** solutions $u(t) \in H^1(\mathbb{R}^n)$ of equation

$$u_t = \Delta u + u^p \quad \text{on } \mathbb{R}^n \times (0, T), \quad 1 < p < \frac{n+2}{n-2}.$$

- u **blows up** in finite time T if $\lim_{t \rightarrow T} \|u(t)\|_{H^1(\mathbb{R}^n)} = +\infty$.
- Recall: $\|u\|_{H^1(\mathbb{R}^n)} := \left(\int_{\mathbb{R}^n} u^2 + |Du|^2 dx \right)^{1/2}$.
- Fillipas, Giga, Herrero, Kohn, Merle, Velazquez, Zaag** etc: analyzed the **blow up** behavior behavior of the solution $u(t)$ near a singular point (a, T) .
- Self-similar scaling:**

$$w(y, s) = (T-t)^{\frac{1}{p-1}} u(x, t), \quad y = \frac{x-a}{\sqrt{T-t}}, \quad s = -\log(T-t).$$

- Giga, Kohn - 1985:** We have $\|w(s)\|_{L^\infty(\mathbb{R}^n)} \leq C, s > -\log T$.

The rescaled semi-linear heat equation

- The **rescaled solution** satisfies the equation

$$(\star) \quad w_s = \Delta w - \frac{1}{2}y \cdot \nabla w - \frac{w}{p-1} + w^p.$$

- To analyze the **blow up** behavior of u we need to understand the **long time** behavior of w as $s \rightarrow +\infty$.
- If $s_k \rightarrow +\infty$ then (passing to a subsequence)

$$w_k(y, s) := w(y, s + s_k)$$

converges to an **eternal** solution W of equation (\star) .

- To understand the blow up behavior of u we need to **classify** the **bounded eternal** solutions of (\star) .

Eternal solutions of the semi-linear heat equation

- Consider a **bounded nonnegative eternal** solutions of equation

$$(\star) \quad w_s = \Delta w - \frac{1}{2}y \cdot \nabla w - \frac{w}{p-1} + w^p \quad \text{on } \mathbb{R}^n \times \mathbb{R}.$$

- Steady states:** $w = 0$ or $w = \kappa$, with $\kappa := (p-1)^{-\frac{1}{p-1}}$.
- $\lim_{s \rightarrow \pm\infty} w(\cdot, s) =$ **steady state**.
- Solutions **independent of y** : $\phi(s) = \kappa(1 + e^s)^{-\frac{1}{p-1}}$.
- Theorem** (Giga-Kohn '87 and Merle-Zaag '98)
If w is bounded nonnegative **eternal** solution of (\star) then

$$w = 0 \quad \text{or} \quad w = \kappa \quad \text{or} \quad w(s) = \phi(s - s_0).$$

- The main **main complexity** of this result is the **classification** of the eternal solutions that connect the steady states $w_{-\infty} = \kappa$ and $w_{+\infty} = 0$. These solutions are shown to be independent of y , hence given by $\phi(s)$.
- Its proof is heavily involved and strongly relies on analyzing the **behavior of the solution w near $\tau \rightarrow -\infty$** in terms of its projections on the positive, zero and negative eigenspaces of the linearized operator $Lv := \Delta v - \frac{1}{2}y \cdot \nabla v + v$ at $w = \kappa$.
- Note that in terms of $v := w - \kappa$, the equation takes the form

$$v_\tau = Lv + f(v)$$

with superlinear error $f(v) := (v + \kappa)^p - \kappa^p - p\kappa^{p-1}v$.

- Similar equations often arise in the analysis of singularities of **neck-pinches** in the mean curvature flow and the Ricci flow.

Ancient solutions to the curve shortening flow

- Let Γ_t be a family of closed curves which is an embedded solution to the **Curve shortening flow**, i.e. the embedding $F : \Gamma_t \rightarrow \mathbb{R}^2$ satisfies

$$\frac{\partial F}{\partial t} = -\kappa \nu$$

with κ the **curvature** of the curve and ν the **outer normal**.

- Gage - 1984, Gage and Hamilton - 1996**: if Γ_0 is convex, then the CSF shrinks Γ_t to a round point.
- Grayson - 1987**: if Γ_0 is any embedded curve in \mathbb{R}^2 , then the solution Γ_t to the CSF does not develop singularities before it becomes strictly convex.
- Problem**: Classify the **ancient compact embedded solutions** to the Curve shortening flow.

The evolution of the curvature

- Let Γ_t is an **ancient convex solution** to the CSF which defined on $I = (-\infty, T)$ and **shrinks to a point** at T .
- The **curvature** κ of Γ_t evolves, in terms of its arc-length s , by

$$\kappa_t = \kappa_{ss} + \kappa^3.$$

- If θ is the *angle between the tangent vector of Γ_t and a fixed direction ω* , then on convex curves $\kappa = \kappa(\theta, t)$ satisfies

$$\kappa_t = \kappa^2 \kappa_{\theta\theta} + \kappa^3.$$

- We introduce the **pressure function** $p = \kappa^2$ which evolves by

$$p_t = p p_{\theta\theta} - \frac{1}{2} p_{\theta}^2 + 2 p^2.$$

Examples and the Result

- **Type I** solution (**contracting circles**): $p(\theta, t) = \frac{1}{2(T-t)}$
- **Type II** solution (**Angenent ovals**): For all $\lambda > 0$ and $\theta_0 \in \mathbb{R}$:

$$p(\theta, t) = \lambda \left(\frac{1}{1 - e^{-2\lambda(T-t)}} - \sin^2(\theta + \theta_0) \right), \quad t < 0$$

As $t \rightarrow -\infty$ the Angenent ovals look like **two grim reaper** solutions to the CSF glued together.

- **Theorem** (D., Hamilton, Sesum)
The **only** ancient convex solutions to the CSF are the contracting spheres or the Angenent ovals.
- **Proof**: It is based on various monotonicity formulas and the fact that as $t \rightarrow T$ any solution becomes circular.

Non-Convex ancient solutions

- **Question:** Do they exist **non convex compact** embedded solutions to the curve shortening flow ?
- **Angenent - 2011:** Presents a **YouTube video** of an ancient solution to the CSF built out from one Yin-Yang spiral and one Grim Reaper.
- It has exactly two inflection points, until it becomes convex.
- **Proof:** at <http://www.youtube.com/watch?v=8Ez0QoJ3XG8>.
- Angenent is currently working on giving a rigorous construction of this !

Open problems - Mean Curvature flow

- **Problem:** Provide the classification of **ancient convex compact** solutions of the **Mean Curvature flow** in dimensions $n \geq 2$.
- **Xu-Jia Wang - 2011:** Classifies the blow-down in space-time as $t \rightarrow -\infty$ of any ancient convex solution to the MCF that sweeps the whole space.
- **Huisken & Sinestrari; Haslhofer & Hershkovits 2013:** Provide geometric conditions that characterize an ancient convex solution to the MCF as a family of contracting spheres.
- **Angenent - 2012:** Provides the **detailed matched asymptotics** of ancient convex compact solutions to MCF that near $t = -\infty$ are (after rescaling) small perturbations of an ellipsoid.
- **B. White - 2003; Haslhofer & Hershkovits - 2013:** Establish the existence of the Angenent ovals.

Ancient solutions of the Ricci flow on S^2

- Consider an **ancient solution** of the **Ricci flow**

$$(RF) \quad \frac{\partial g_{ij}}{\partial t} = -2 R_{ij}$$

on S^2 that exists for all time $-\infty < t < T$ and becomes singular at time T .

- In dim 2, we have $R_{ij} = \frac{1}{2}R g_{ij}$, where R is the scalar curvature.
- **R. Hamilton 1988, Chow 1991**: After re-normalization, the metric becomes **spherical** at $t = T$.
- **Problem**: Classify the ancient compact solutions.

The equations for the conformal factor and the pressure

- Choose a parametrization $g_{S^2} = d\psi^2 + \cos^2 \psi d\theta^2$ of the limiting spherical metric and parametrize the (RF) by this, i.e. we write $g(\cdot, t) = u(\cdot, t) g_{S^2}$.
- If $g(\cdot, t) = u(\cdot, t) g_{S^2}$, then the (RF) becomes equivalent to:

$$u_t = \Delta_{S^2} \log u - 2, \quad \text{on } S^2 \times (-\infty, T).$$

- Assume from now on that $T = 0$.
- It is natural to consider the **pressure** function $v = u^{-1}$ which evolves by

$$(PE) \quad v_t = v \Delta_{S^2} v - |\nabla v|^2 + 2v^2.$$

Examples of ancient solutions on S^2

- Example of a **Type I** solution (**contracting spheres**):

$$v(\psi, \theta, t) = \frac{1}{2(-t)}$$

- Example of a **Type II** solution is the **King** solution:

$$v(\psi, \theta, t) = -\mu [\coth(2\mu t) - \tanh(2\mu t) \sin^2 \psi]$$

for $t < 0$, with $\mu > 0$.

- As $t \rightarrow -\infty$ the King solution looks like two **cigar solutions** glued together.

The King solutions

- We look for explicit **ancient** solutions $g_{ij} = u(\cdot, t) g_{S^2}$ of the (RF) that are different than the contracting spheres and become singular at time $T = 0$.
- It is simpler to work on \mathbb{R}^2 (via stereographic projection) setting $g_{ij} = \bar{u} g_{euc}$.
- The function \bar{u} satisfies the logarithmic **fast-diffusion** equation

$$\text{(LFD)} \quad \bar{u}_t = \Delta \log \bar{u} \quad \text{on } \mathbb{R}^2 \times (-\infty, 0).$$

- **J.R. King - 1993** looked for radial solutions of equation (LFD) where the **pressure** function $\bar{v} := \bar{u}^{-1}$ is a **polynomial** function in $r := |x|$ with coefficients depending on t .

The King solutions

- A direct calculation shows that

$$\bar{v}(x, t) = a(t) + 2b(t)|x|^2 + a(t)|x|^4$$

where either $a(t) = b(t)$ (recovering the [contracting spheres](#))
or

$$a(t) = -\frac{\mu}{2} \operatorname{csch}(4\mu t), \quad a(t) = -\frac{\mu}{2} \operatorname{coth}(4\mu t)$$

for any $\mu > 0$.

- These solutions were independently discovered by P. Rosenau.
- The King solutions are **not self-similar**.
- As $t \rightarrow -\infty$ they look like two [cigar \(Barenblatt self similar\)](#) solutions glued together.

The classification result

Theorem: (D., Hamilton, Sesum)

An **ancient** solution to the (RF) on S^2 is either one of the **contracting spheres** or one of the **King** solutions.

Sketch of proof:

We recall that if $g(\cdot, t) = u(\cdot, t) g_{S^2}$ is the evolving metric, then the *pressure* $v = u^{-1}$ satisfies:

$$v_t = v \Delta v - |\nabla v|^2 + 2v^2 = R v > 0$$

- We show, by establishing sharp a priori derivative estimates, that $v(\cdot, t) \xrightarrow{C^{1,\alpha}} v_\infty$, as $t \rightarrow -\infty$, for all $\alpha < 1$.
- Via a suitable **Lyapunov functional** we show that

$$R_\infty := \lim_{t \rightarrow -\infty} R(\cdot, t) = 0 \quad \text{a.e. on } S^2$$

and

$$v_\infty \Delta v_\infty - |\nabla v_\infty|^2 + 2v_\infty^2 = 0 \quad \text{a.e. on } S^2$$

Sketch of proof - Continuation

- We next classify the steady states v_∞ which satisfy

$$v_\infty \Delta v_\infty - |\nabla v_\infty|^2 + 2v_\infty^2 = 0.$$

Main Step: We show that v_∞ has **at most two zeros**.

We conclude that

$$v_\infty(\psi, \theta) = C \cos^2 \psi, \quad \text{for } C \geq 0$$

namely that the pointwise limit as $t \rightarrow -\infty$ is a **cylinder**.

- If $C = 0$, then v must be a family of **contracting spheres**.
- If $C > 0$, then v must be one of the **King solutions**.

The characterization of King solutions

- To capture the **King** solutions we consider the scaling invariant **nonotone quantity**

$$Q(x, y, t) := \bar{v} [(\bar{v}_{xxx} - 3\bar{v}_{xyy})^2 + (\bar{v}_{yyy} - 3\bar{v}_{xxy})^2]$$

where $\bar{v} := \bar{u}^{-1}$ is the pressure in plane coordinates.

- Using complex variable notation $z = x + iy$, this quantity is nothing but

$$Q = \bar{v} |\bar{v}_{zzz}|^2.$$

- The quantity Q is well defined.
- It turns out that $Q \equiv 0$ implies that \bar{v} is one of the **King** solutions.
- To establish that $Q \equiv 0$ we prove that:
 - $Q_{\max}(t)$ is decreasing in t (by considering its evolution equation), and
 - $\lim_{t \rightarrow -\infty} Q_{\max}(t) = 0$.

The 3 dimensional Ricci flow - Open problems

- **3-dim Ricci flow:** The equivalent to the King solutions have been shown to exist by [Perelman](#). These solutions are Type II ancient solutions and they are **k non-collapsing**.
- Other compact solutions in closed form have been found by [V.A. Fateev](#) in a paper dated back to 1996.
- **Conjecture:** The only **k non-collapsing** ancient compact solutions to the 3-dim Ricci flow are the contracting spheres and the [Perelman](#) solutions. In particular they are radially symmetric.
- [Brendle, Huisken, Sinestrari - 2011](#): Provided a **pinching curvature condition** that characterizes ancient solutions to the 3-dim Ricci as contracting spheres.

Ancient solutions to the Yamabe flow

- Consider an **ancient solution** g_{ij} of the **Yamabe flow** ($n \geq 3$)

$$\frac{\partial g_{ij}}{\partial t} = -R g_{ij} \quad \text{on } -\infty < t < T$$

where the evolving metric g_{ij} is conformally equivalent to the standard metric on S^n , namely $g_{ij} = u^{\frac{4}{n+2}} g_{S^n}$.

- **Question:** Can we classify all such ancient solutions ?

The Yamabe flow - Background

- Let (M, g_0) be a compact manifold without boundary of dimension $n \geq 3$. If $g = v^{\frac{4}{n-2}} g_0$ is a metric conformal to g_0 , the **scalar curvature** R of g is given in terms of v and the scalar curvature R_0 of g_0 by

$$R = -v^{-\frac{n+2}{n-2}} (\bar{c}_n \Delta_{g_0} v - R_0 v).$$

- In 1989 R. Hamilton introduced the **Yamabe flow**

$$\frac{\partial g_{ij}}{\partial t} = -R g_{ij}.$$

as a parabolic approach to solve the **Yamabe problem**.

- Brendle - 2007**: convergence of the normalized flow to a metric of constant scalar curvature (under some mild technical assumptions in dimensions $n \geq 6$).
- Previous important works: **Hamilton '89**, **Chow '92**, **Ye '94**, **del Pino-Saez '2001**, **Schwetlick-Struwe '2003**.

Ancient compact solutions to the Yamabe flow

- **King '93:** found non-self similar ancient compact solutions to the Yamabe flow in closed form. As $t \rightarrow -\infty$ they converge to two **cigar solutions** joined with a long cylindrical neck.
- They resemble the King Ricci flow solutions, but they are of **type I**.
- **Question 1:**
Are the **contracting spheres** and the **King solutions** the only **Type I** ancient solutions of the Yamabe flow (YF) that are conformal to the standard spherical metric on S^n ?
- **Question 2:**
Are there any **Type II** ancient solutions ?

Towers of bubbles

- The following is joint work with [M. del Pino](#) and [N. Sesum](#).
- We construct a class of rotationally symmetric **ancient solutions**

$$g_{ij} = u^{\frac{n+2}{4}}(\cdot, t) g_{S^n}$$

of the **Yamabe flow** on $-\infty < t < 0$ that become singular at $t = 0$ and near $t = -\infty$ they look (after re-normalization) like **n spheres** jointed by **short necks**.

- The curvature operator of these solutions **changes sign** near $t = -\infty$ and they are **Type II** ancient solutions.
- We refer to them as **towers of bubbles**.
- Our construction can be viewed as a parabolic analogue of the **gluing** of manifolds of constant scalar curvature or constant mean curvature that has been used in the past in various **elliptic settings**.

The equations

- **Notation:** From now on we will set $m = \frac{n-2}{n+2}$ and $p = \frac{n+2}{n-2}$.
- Express, for the moment, the metric g_{ij} as $g_{ij} = \bar{u}^{\frac{n+2}{4}} g_{euc}$, where g_{euc} denotes the standard Euclidean metric.
- Our problem is equivalent to constructing a radially symmetric solution $\bar{u}(|x|, t)$ of the **fast-diffusion** equation

$$(FD) \quad \bar{u}_t = \Delta \bar{u}^m \quad \text{on } \mathbb{R}^n \times (-\infty, 0)$$

with the right asymptotic behavior as $r = |x| \rightarrow \infty$.

- The function $\hat{u} := u^m$ satisfies the equation

$$\hat{u}_t^p = \Delta \hat{u} \quad \text{on } \mathbb{R}^n \times (-\infty, 0)$$

The equation in cylindrical coordinates

- We next introduce **cylindrical coordinates** setting

$$v(x, \tau) = (-t)^{-\frac{1}{p-1}} r^{\frac{2}{p-1}} \hat{u}(r, t), \quad r = e^x, \quad t = 1 - e^{-\tau}$$

- In these coordinates the equation becomes

$$(v^p)_\tau = v_{xx} + \alpha v^p - \beta v$$

for certain constants $\alpha > 0$ and $\beta > 0$, depending on dimension n . It is defined for all $(x, \tau) \in \mathbb{R} \times \mathbb{R}$.

- By scaling away the constants, we may assume that $\alpha = \beta = 1$.

The equation in cylindrical coordinates

- Consider from now on our equation in cylindrical coordinates

$$(E) \quad (u^p)_t = u_{xx} - u + u^p, \quad \text{on } \mathbb{R} \times (-\infty, +\infty)$$

for some $t_0 \in \mathbb{R}$ and $p = \frac{n+2}{n-2}$.

- Notation:** For simplicity we denote τ by t .

- The function $w(x) = \left(\frac{2k_n}{e^{\beta x} + e^{-\beta x}} \right)^{1/\beta}$ with $\beta = \frac{2}{n-2}$ and $k_n = \left(\frac{n}{n-2} \right)^{1/2}$ solves the steady-state equation

$$w_{xx} - w + w^p = 0, \quad x \in \mathbb{R}$$

and represents the conformal factor for the standard metric on the unit sphere in cylindrical coordinates.

The ansatz of the non-linear problem

- We construct an ancient solution of (E) of the form

$$u(x, t) = (1 + \eta(t)) z(x, t) + \psi(x, t)$$

for a suitable small **parameter** function $\eta(t)$, where

$$z(x, t) = w(x - \xi(t)) + w(x + \xi(t))$$

and $\psi(x, t) \rightarrow 0$ as $t \rightarrow -\infty$ in a suitable sense.

- The function $\xi(t)$ is given by

$$\xi(t) = \xi_0(t) + h(t), \quad \xi_0(t) \approx \frac{1}{2} \log(2b|t|)$$

for a suitable $b > 0$ and a small **parameter** function $h(t)$.

- Both parameter functions $h(t)$ and $\eta(t)$ will **decay** in $|t|$, as $|t| \rightarrow -\infty$ in a suitable sense.

The ansatz of the non-linear problem

- We notice that $z(\cdot, t)$ is an **even function** of x and we will impose the condition that $\psi(\cdot, t)$ is an even function as well.
- A direct computation shows that equation (E) for u is equivalent to equation

$$(\star_{NL}) \quad pz^{p-1} \psi_t = \psi_{xx} - \psi + pz^{p-1} \psi + z^{p-1} E(\psi, h, \eta)$$

for the unknown **perturbation** function $\psi(x, t)$.

- Our goal is to construct an ancient solution ψ of the nonlinear equation (\star_{NL}) via the **fixed point Theorem**.
- To this purpose we will first need to **invert** the corresponding linear problem for any given **parameter functions** h and η in suitable spaces.

The linear problem

- We first establish the existence of an **ancient** solution ψ to

$$(\star_L) \quad \rho z^{\rho-1} \psi_t = \psi_{xx} - \psi + \rho z^{\rho-1} \psi + z^{\rho-1} g$$

in **suitable Banach spaces**.

- Recall that $z(x, t) = w(x - \xi(t)) + w(x + \xi(t))$.
- We will consider a class of even functions g that **decay** both in x and t at suitable rates and satisfy certain **orthogonality conditions**.
- We build an even **ancient** solution $\psi := T(g)$ of equation (\star_L) such that

$$\|\psi\|_{\star\star} \leq C \|g\|_{\star}$$

and shares the same **decay rates** and **orthogonality conditions**.

- $\|\cdot\|_{\star\star}$ in an appropriate weighted L^2 and $W^{2,p}$ norm.

The orthogonality conditions

- Recall our equation

$$(\star_L) \quad \rho z^{p-1} \psi_t = \psi_{xx} - \psi + \rho z^{p-1} \psi + z^{p-1} g$$

on $\mathbb{R} \times (-\infty, t_0]$, for some $t_0 \in \mathbb{R}$.

- We say that $g \in \bar{\mathcal{S}}$, if g satisfies the **orthogonality conditions**

$$\int_{\mathbb{R}} g(y \pm \xi(t), t) w'(y) w^{p-1} dy = 0, \quad \text{a.e. } t < t_0$$

$$\int_{\mathbb{R}} g(y \pm \xi(t), t) w(y) w^{p-1} dy = 0, \quad \text{a.e. } t < t_0.$$

- Notice that since g is assumed to be an **even function** in x , the above orthogonality conditions also imply the symmetric conditions $f(y, t) := g(y + \xi(t)) \in \mathcal{S}$.

Back to the non-linear problem

- Our goal is to establish the existence of an ancient solution of the non-linear equation

$$(\star_{NL}) \quad \rho z^{p-1} \psi_t = \psi_{xx} - \psi + \rho z^{p-1} \psi + z^{p-1} E(\psi, h, \eta)$$

on $\mathbb{R} \times (-\infty, t_0]$, for t_0 sufficiently close to $-\infty$, such that $\psi(\cdot, t) \rightarrow 0$, as $t \rightarrow -\infty$ in a suitable sense.

- This is equivalent to $u := (1 + \eta) z(x, t) + \psi(x, t)$, with

$$z(x, t) = w(x - \xi) + w(x + \xi), \quad \xi = \xi_0 + h$$

being the desired ancient solution of the Yamabe flow.

- We will see that $\xi_0(t) \approx \frac{1}{2} \log(2b|t|)$, as $t \rightarrow -\infty$.

The auxiliary problem

- We will establish the existence of ψ through an iteration process, where at each step ψ_{k+1} we will be the solution of the linear equation with right hand side $E(\psi_k, h, \eta)$.
- Hence it is essential that $\psi_k, E(\psi_k, h, \eta) \in \bar{\mathcal{S}}$, for all k .
- Because of this, we will first consider the **auxiliary system**

$$(\star_A) \quad pz^{p-1} \psi_t = \psi_{xx} - \psi + pz^{p-1} \psi + z^{p-1} \bar{E}(\psi, h, \eta)$$

where $\bar{E}(\psi, h, \eta) := E(\psi, h, \eta) - (c_1 z + c_2 \bar{z})$ and $c_1 := c_1(t)$, $c_2 = c_2(t)$ are chosen so that

$$\bar{E}(\psi, h, \eta) \in \bar{\mathcal{S}}, \quad \text{if } \psi \in \bar{\mathcal{S}}$$

Recall $z = w(x - \xi) + w(x + \xi)$, $\bar{z} = w'(x - \xi) + w'(x + \xi)$.

Solving the nonlinear problem

- For any given (h, η) , we establish the existence of a solution $\psi = \Psi(h, \eta)$ of the auxiliary equation where $c_1(t)$ and $c_2(t)$ are chosen so that

$$\bar{E}(\psi, h, \eta) := E(\psi, h, \eta) - (c_1(t)z + c_2(t)\bar{z})$$

satisfies our orthogonality conditions whenever ψ does.

- Thus $\psi = \Psi(h, \eta)$ will define a solution to our original equation if we manage to adjust the parameter functions (h, η) in such a way that $c_1 \equiv 0$ and $c_2 \equiv 0$.
- This is equivalent to choosing (h, η) so that

$$\int_{\mathbb{R}} E(\psi) w^p(x + \xi(t)) dx = 0$$

and

$$\int_{\mathbb{R}} E(\psi) w'(x + \xi(t)) w^{p-1}(x + \xi(t)) dx = 0.$$

Solving for h and η

- We conclude that (ξ, η) with $\xi = \xi_0 + h$ satisfy the system

$$(\star\star_1) \quad \dot{\eta} - \frac{p-1}{p}\eta - \alpha e^{-2\xi} = \mathcal{R}_1(\xi, \dot{\xi}, \eta, \dot{\eta})$$

$$(\star\star_2) \quad \dot{\xi} + \beta e^{-2\xi} = \mathcal{R}_2(\xi, \dot{\xi}, \eta, \dot{\eta}).$$

- We observe that the ODE $\dot{\xi} + \beta e^{-2\xi} = 0$ admits the explicit solutions

$$\xi_0(t) := \frac{1}{2} \log(2b|t - \lambda|), \quad \text{for any } \lambda \in \mathbb{R}$$

where λ is a translating parameter (dilation in the original variables).

- We conclude the existence of $\xi = \xi_0 + h$ and η via a fixed point Theorem.

The Main Result

- The previous discussion leads to the final result:

Theorem. (*D., M. del Pino, N. Sesum*)

There exist radially symmetric ancient solutions $g_{ij} = u g_{cyl}$ of the Yamabe flow of the form

$$u(x, t) = (1 + \eta) [w(x - \xi) + w(x + \xi)] + \psi(x, t)$$

with

$$\xi := \frac{1}{2} \log(2b|t|) + h(t)$$

for suitable parameter functions h and η that **decay** in $|t|$ as $t \rightarrow -\infty$ at a certain rate. This ancient solution is of **Type II** and its curvature operator **changes sign**.

- **Remark:** The result in the previous theorem generalizes to any finite number of k -bubbles.

- The appearance of the towers of bubbles indicates that the problem of the **classification** of all ancient solutions to the Yamabe flow on S^n is delicate if not impossible.
- **Question:** What are the **Type I** ancient solutions ?
- **Question:** Are the contracting spheres and the King solutions the only type I ancient solutions ?
- **Question:** Are all Type I ancient solutions rotationally symmetric ?

Predictions - The new King solutions

- **Question:** Are the contracting spheres and the King solutions the only type I ancient solutions ?
- **D., King and Sesum:** There exist a two parameter family of self-similar solutions $g_\gamma := v_{\gamma,\lambda}^{\frac{4}{n+2}} dx^2$ to the Yamabe flow on \mathbb{R}^n that they are complete, non-compact and all behave as **cylinders** as $|x| \rightarrow \infty$.
- $\gamma = 2$ corresponds to the explicit Barenblatt solutions.
- **Conjecture:** One may construct new ancient King solutions such that as $t \rightarrow -\infty$ look as two g_γ solutions joint with a neck.