# Small scale formations for the incompressible Boussinesq equation

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## 2D Boussinesq equation without density diffusivity

- $\rho(x, t)$ : density of incompressible fluid.
- u(x, t): velocity field of fluid.
- The spatial domain  $\Omega$  is either the plane  $\mathbb{R}^2$ , the torus  $\mathbb{T}^2$ , or the strip  $\mathbb{T} \times [-\pi, \pi]$ .
- Throughout this talk, we consider the 2D Boussinesq equation without density diffusivity:

$$\begin{cases} \rho_t + u \cdot \nabla \rho = \mathbf{0}, \\ u_t + u \cdot \nabla u = -\nabla p - \rho e_2 + \nu \Delta u, \\ \nabla \cdot u = \mathbf{0}, \end{cases}$$



- We'll discuss the viscous case  $\nu > 0$ , and the inviscid case  $\nu = 0$ .
- Goal: In both cases, we'll prove that solution can have small scale formation (infinite-in-time growth of Sobolev norms) as  $t \to \infty$ .

The viscous case: global well-posedness and upper bounds

When  $\nu > 0$ , global-wellposedness of regular solutions is known:

- For  $\Omega = \mathbb{R}^2$ : global regularity by Hou–Li '05 in the space  $(u, \rho) \in H^m \times H^{m-1}$  for  $m \ge 3$ , and Chae '06 in  $H^m \times H^m$  for  $m \ge 3$ .
- For bounded  $\Omega$ : global regularity by Lan–Pan–Zhao '11 in  $H^3 \times H^3$ , and Hu–Kukavica–Ziane '13 in  $H^m \times H^{m-1}$  for  $m \ge 2$ .

Upper bounds for the global solution:

- Ju '17: For bounded  $\Omega$ ,  $\|\rho\|_{H^1} \lesssim e^{Ct^2}$ ;
- Kukavica–Wang '20: For bounded  $\Omega$ ,  $\|\rho\|_{H^1} \leq e^{Ct}$  and  $\|u\|_{W^{2,p}} \leq C_p$ ; for  $\mathbb{R}^2$ ,  $\|\rho\|_{H^1} \leq e^{Ct^{(1+\beta)}}$ .
- Kukavica-Massatt-Ziane '21: For bounded  $\Omega$ ,  $\|\rho\|_{H^2} \leq C_{\epsilon} e^{\epsilon t}, \|u\|_{H^3} \leq C_{\epsilon} e^{\epsilon t}.$

- Note that the above estimates all deal with the upper bounds of solutions.
- Question. What about lower bounds? Can solutions have small scale formation as  $t \to \infty$ ?
- Lower bound by Brandolese–Schonbek '12: in  $\mathbb{R}^2$ , if  $\rho_0$  does not have mean zero,  $\|u(t)\|_{L^2} \sim (1+t)^{1/4}$ . (This is due to potential energy converting to kinetic energy, and does not imply growth in higher derivatives)
- We are not aware of any examples of infinite-in-time growth of  $\|\rho(t)\|_{H^m}$  in the literature.

## Small scale formation in the viscous case

#### Theorem (Kiselev–Park–Y. '22, preprint)

Let  $\nu > 0$ ,  $\Omega = \mathbb{T}^2$ . If the smooth initial data  $(\rho_0, u_0)$  satisfies the following

• Symmetry assumptions:  $\rho_0$  is even-odd,  $u_{01}$  is odd-even,  $u_{02}$  is even-odd

• Sign assumptions:  $\rho_0 \ge 0$  for  $x_2 \ge 0$ , and  $\rho_0 = 0$  on the  $x_2$ -axis.

Then the global-in-time smooth solution satisfies

$$\limsup_{t\to\infty} \frac{t^{-\frac{1}{6}}}{\|\rho(t)\|_{\dot{H}^1}} = +\infty.$$

Preserved for all time!

Remark: Under these assumptions one can show  $\|\rho(t)\|_{\dot{H}^1}$  has a refined sub-exponential upper bound  $\exp(Ct^{\alpha})$  for some  $\alpha \in (0, 1)$ , so the growth is somewhere between algebraic and sub-exponential.



# Evolution of potential energy

- Define the potential energy  $E_P(t) := \int_{\mathbb{T}^2} \rho x_2 dx$ , and kinetic energy  $E_K(t) := \int_{\mathbb{T}^2} |u|^2 dx$ .
- It's well-known that the total energy is decreasing in time:

$$\frac{d}{dt}(E_P(t)+E_K(t))=-\nu\|\nabla u(t)\|_{L^2}^2.$$

This implies that  $\int_0^\infty \|\nabla u(t)\|_{L^2}^2 dt < C(\nu, \rho_0, u_0).$ 

- Since the two equations are coupled by the gravity force, we'll track the evolution of potential energy  $E_P(t)$  itself.
- A quick computation gives  $\frac{d}{dt}E_P(t) = \int_{\mathbb{T}^2} \rho u_2 dx$ , which is uniformly bounded.
- Let's take one more time derivative:

$$\frac{d^2}{dt^2}E_P(t) = \underbrace{\sum_{i,j=1}^2 \int_{\mathbb{T}^2} ((-\Delta)^{-1}\partial_2 \rho) \partial_i u_j \partial_j u_i dx}_{=:A(t)} - \underbrace{\nu \int_{\mathbb{T}^2} \nabla \rho \cdot \nabla u_2 dx}_{=:B(t)} - \frac{\|\partial_1 \rho\|_{\dot{H}^{-1}}^2}{=:B(t)}$$

- Since  $\int_0^\infty \|\nabla u(t)\|_{L^2}^2 dt < \infty$ , this implies  $\int_0^\infty A(t) dt < \infty$ .
- Suppose  $\limsup_{t\to\infty} \|\nabla\rho\|_{L^2} < \infty$ , we have  $\int_0^\infty B(t) dt \lesssim t^{1/2}$ .
- This implies  $\int_0^\infty \|\partial_1 \rho\|_{\dot{H}^{-1}}^2 dt \lesssim t^{1/2}$ , so  $\|\partial_1 \rho\|_{\dot{H}^{-1}}^2$  needs to decay to zero like  $t^{-1/2}$  as  $t \to \infty$ .
- Key observation (by a Fourier argument): If  $\|\partial_1 \rho(t)\|_{\dot{H}^{-1}} \ll 1$  and  $\rho \equiv 0$  on  $x_2$  axis, we have  $\|\rho\|_{\dot{H}^1} \gg 1$ . More precisely,  $\|\rho\|_{\dot{H}^1} \gtrsim \|\partial_1 \rho(t)\|_{\dot{H}^{-1}}^{-1}$ .



• This contradicts our assumption  $\limsup_{t\to\infty} \|\nabla\rho\|_{L^2} < \infty$ . (A more careful argument gives us algebraic growth in time).

### Inviscid 2D Boussinesq equation

• In the inviscid case  $\mu = 0$ , let us work with the variables  $\rho$  and vorticity  $\omega$ :

$$\begin{cases} \rho_t + \mathbf{u} \cdot \nabla \rho = \mathbf{0}, \\ \omega_t + \mathbf{u} \cdot \nabla \omega = -\partial_1 \rho, \end{cases}$$

where u can be recovered from the Biot-Savart law  $u = \nabla^{\perp} (-\Delta)^{-1} \omega$ .

- Whether smooth initial data can lead to a blow-up in  $\mathbb{T}^2$  or  $\mathbb{R}^2$  is an outstanding open question.
- It is well-known that away from the axis of symmetry, the 3D axisymmetric Euler equation is closely related to 2D Boussinesq:

$$\begin{cases} D_t(ru^{\theta}) = 0, \\ D_t\left(\frac{\omega^{\theta}}{r}\right) = r^{-4}\partial_z(ru^{\theta})^2, \end{cases}$$

where  $D_t := \partial_t + u^r \partial_r + u^z \partial_z$  is the material derivative, and  $(u^r, u^z)$  can be recovered from  $\omega^{\theta}/r$  by a similar Biot-Savart law.

In the presence of boundary, or for non-smooth initial data, there are many exciting developments on finite-time blow-up:

- Luo-Hou '14: convincing numerical evidence for blow-up at the boundary for 3D axisymmetric Euler
- Elgindi–Jeong '20: blow-up in domain with a corner
- Elgindi '21: blow-up for  $C^{1,\alpha}$  solutions for 3D Euler
- Chen-Hou '21: blow-up for  $C^{1,\alpha}$  solutions with boundary
- Wang–Lai–Gómez-Serrano–Buckmaster '22: numerics for approximate self-similar blow-up solution using physics-informed neural networks.
- Chen–Hou '22: stable nearly self-similar blowup for smooth solutions (combination of analysis + computer-assisted estimates)

**Question:** Can one construct solutions with infinite-in-time growth for more general class of initial data?

#### Theorem (Kiselev–Park–Y. '22, preprint)

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Let  $\Omega = \mathbb{T} \times [0, \pi]$ . Let  $\rho_0 \in C^{\infty}(\Omega)$  be even in  $x_1$ , and  $\omega_0 \in C^{\infty}(\Omega)$  be odd in  $x_1$ , with  $\int_{[0,\pi]\times[0,\pi]} \omega_0 dx \ge 0$ . Assume that there exists  $k_0 > 0$  such that  $\rho_0 \ge k_0 > 0$  on  $\{0\} \times [0,\pi]$ , and  $\rho_0 \le 0$  on  $\{\pi\} \times [0,\pi]$ . Then the solution satisfies the following during its lifespan:

$$egin{aligned} \|\omega(t)\|_{L^p(\Omega)} \gtrsim t^{3-rac{2}{p}}, \ \|u(t)\|_{L^\infty(\Omega)} \gtrsim t, \ \sup_{\in [0,t]} \|
abla 
ho( au)\|_{L^\infty(\Omega)} \gtrsim t^2 \end{aligned}$$



The proof is a soft argument, based on an interplay between various monotone and conservative quantities.

## Monotonicity of vorticity integral

• Let Q be the right half of the strip. Simple but useful observation:

$$\int_{Q} \frac{d}{dt} \int_{Q} \omega dx = \int_{Q} \frac{1}{\sqrt{\omega dx}} \int_{Q} \frac{\partial_{1} \rho dx}{\partial t}$$
$$= \int_{0}^{\pi} \frac{\rho(0, x_{2}, t)}{\sum k_{0}} dx_{2} - \int_{0}^{\pi} \frac{\rho(\pi, x_{2}, t)}{\sum k_{0}} dx_{2}$$
$$\geq k_{0}\pi.$$

• Since  $\int_{\partial Q} u \cdot dl = \int_{Q} \omega dx \ge k_0 \pi t$ , we have  $\|u(t)\|_{L^{\infty}}$  grows at least linearly.

- On the other hand,  $||u||_{L^2}$  is bounded for all times by energy conservation.
- Combining the boundedness of ||u||<sub>L<sup>2</sup>(Q)</sub> and linear growth of ∫<sub>∂Q</sub> u · dl, we know u must change rapidly in a small neighborhood of ∂Q, leading to super-linear growth of ∇u (and ω).

# Infinite-in-time growth in $\mathbb{T}^2$

- To our best knowledge, there has been no blow-up / infinite-in-time growth results in  $\mathbb{T}^2.$
- In T<sup>2</sup>, we obtain infinite-in-time growth for a large class of initial data satisfying certain symmetry/sign conditions:

Theorem (Kiselev–Park–Y. '22, preprint)

Let  $\rho_0 \in C^{\infty}(\mathbb{T}^2)$  be even-odd, and  $\omega_0 \in C^{\infty}(\mathbb{T}^2)$  be odd-odd. Assume  $\rho_0 \ge 0$ on  $\{0\} \times [0, \pi]$  with  $k_0 := \sup_{x_2 \in [0, \pi]} \rho_0(0, x_2) > 0$ , and  $\rho_0 \le 0$  on  $\{\pi\} \times [0, \pi]$ . Then the solution satisfies the following during its lifespan:

$$\sup_{\tau\in[0,t]} \|\nabla\rho(\tau)\|_{L^{\infty}(\mathbb{T}^2)} \gtrsim t^{1/2}.$$
 (1)



## 3D axisymmetric Euler in an annular cylinder

Using a similar idea, we obtain infinite-in-time growth for the 3D axisymmetric Euler equation in an annular cylinder

$$\Omega = \{(r, heta, z) : r \in [\pi, 2\pi]; heta \in \mathbb{T}, z \in \mathbb{T}\}.$$

### Theorem (Kiselev–Park–Y. '22, preprint)

Let  $u_0^{\theta} \in C^{\infty}(\Omega)$  be even in z,  $\omega_0^{\theta} \in C^{\infty}(\Omega)$  odd in z, with  $\int_0^{\pi} \int_{\pi}^{2\pi} \omega_0^{\theta} dr dz \ge 0$ . Assume there exists  $k_0 > 0$  such that  $u_0^{\theta} \ge k_0 > 0$  on  $z = \pi$ , and  $|u_0^{\theta}| \le \frac{1}{8}k_0$  on z = 0. Then the solution to axisymmetric 3D Euler satisfies

$$\|\omega^{ heta}(t)\|_{L^p(\Omega)}\gtrsim t^{3-rac{2}{p}} \quad ext{ and } \|u(t)\|_{L^\infty(\Omega)}\gtrsim t$$

during the lifespan of the solution.



## Thank you for your attention!

